

$$\frac{1}{\tau_n} = \frac{2\pi}{\hbar} |W_{mn}|^2 D(E_m) \delta(E_m - E_n)$$

Example

Considers a harmonic perturbation:

$$\hat{W}(t > 0) = \hat{a} e^{i\omega t} + \hat{a}^\dagger e^{-i\omega t}$$

Then the transition rate of scattering from state $|n\rangle$ to $|m\rangle$ is:

$$\frac{1}{\tau_n} = \frac{2\pi}{\hbar} |W_{mn}|^2 \delta(E_m - E_n) = \frac{2\pi}{\hbar} |\langle m | \hat{a} e^{i\omega t} + \hat{a}^\dagger e^{-i\omega t} | n \rangle|^2 \delta(E_m - E_n)$$

$$\frac{1}{\tau_n} = \frac{2\pi}{\hbar} |\langle m | \hat{a} | n \rangle|^2 \delta(E_m - E_n + \hbar\omega) + \frac{2\pi}{\hbar} |\langle m | \hat{a}^\dagger | n \rangle|^2 \delta(E_m - E_n - \hbar\omega)$$

Final energy $E_m = E_n - \hbar\omega$

$E_m = E_n + \hbar\omega$

Ex./ stimulated emission
of a photon

Ex./ photon absorption

Ex./ Phonon emission

Ex./ Phonon absorption

If there are continuous energy states for the $|m\rangle$ (final state):

$$\frac{1}{\tau_n}{}^{emi} = \frac{2\pi}{\hbar} |\langle m | \hat{a} | n \rangle|^2 D(E_m) \delta(E_m - E_n + \hbar\omega)$$

$$\frac{1}{\tau_n}{}^{abs} = \frac{2\pi}{\hbar} |\langle m | \hat{a}^\dagger | n \rangle|^2 D(E_m) \delta(E_m - E_n - \hbar\omega)$$

Since W_{mn} is hermitian (physical):

$$\langle m | \hat{a} | n \rangle = \langle m | \hat{a} | n \rangle^\dagger = \langle n | \hat{a}^\dagger | m \rangle$$

$$\Rightarrow \tau_n{}^{emi} D(E_m = E_n - \hbar\omega) = \tau_n{}^{abs} D(E_m = E_n + \hbar\omega)$$

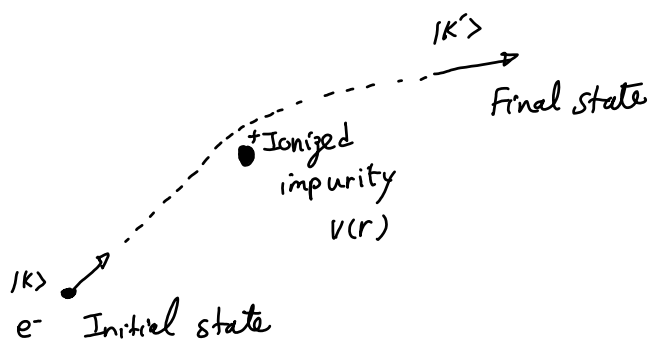
What approximations we used

to drive the Fermi Golden Rule?

- ① $t \rightarrow \infty$ limit So the collision must be complete before the next collision happens.
- ② We assumed the perturbation from $|n\rangle$ is so small that $|a_n|^2 = 1$ (conservation of number of particles ignored)
- ③ Also, we sometime assume that the incident wave is planar $e^{ikx - i\omega t}$ and the final state also is planar $e^{ik'x - i\omega t}$. This means assuming that the collision is localized in real space.

All of these assumptions can be violated in modern semiconductor devices. So use FGR with caution!

Ionized Impurities scattering



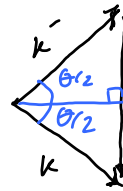
$$\frac{1}{\tau(k, k')} = \frac{2\pi}{\hbar} |W_{kk'}|^2 D(E_{k'}) \delta(E_{k'} - E_k)$$

Must calculate the matrix element $W_{kk'}$:

$$W_{kk'} = \langle k' | v(r) | k \rangle = \int \varphi^*(k') v(r) \varphi(k) d^3r$$

Assume plane wave:
(free electron wavefunction)

$$W_{kk'} = \int e^{-i(\mathbf{k}-\mathbf{k}') \cdot \mathbf{r}} v(r) d^3r$$



$$|\mathbf{k}| = |\mathbf{k}'|$$

$$|q| = 2|\mathbf{k}| \sin \frac{\theta}{2}$$

$$= \int e^{-i\vec{q} \cdot \vec{r}} v(r) d^3r = v(q) \quad \text{Fourier transform}$$

So with plane wave assumption, the matrix element

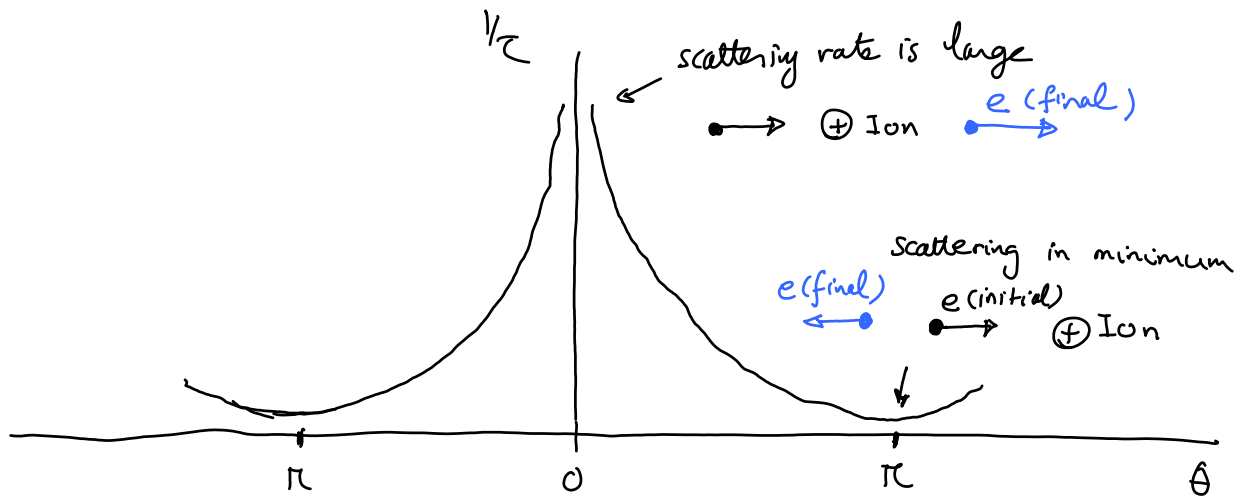
for the transition is the Fourier transform of the
perturbing potential:

$$W_{kk'} = v(q) \quad \vec{q} \equiv \vec{k} - \vec{k}'$$

This is a general argument for
any potential with elastic scattering.

$$v(r) = \frac{-e^2}{4\pi\epsilon_r\epsilon_0} \frac{1}{r} \rightarrow v(q) = \int d^3r v(r) e^{-i\vec{q} \cdot \vec{r}} = \frac{-e^2}{4\pi\epsilon_r\epsilon_0} \int d^3r \frac{e^{-i\vec{q} \cdot \vec{r}}}{r}$$

$$= \frac{-e^2}{4\pi\epsilon_r\epsilon_0} \int_0^\infty dr \int_0^\pi d\theta \int_0^{2\pi} d\varphi \frac{e^{-iqr \cos\theta}}{r} r^2 \sin\theta$$



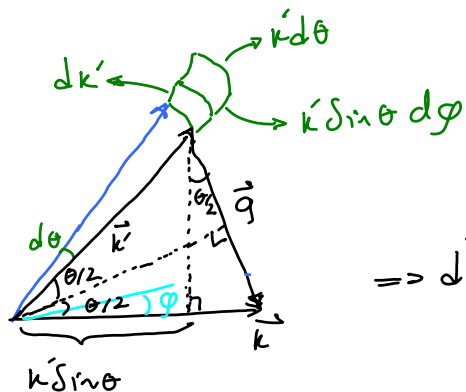
So most electrons are scattered without a significant change in their direction.

The total scattering rate is the integral of $\frac{1}{\tau(k, k')}$ over

all possible final states k' :

$$\frac{1}{\tau(k)} = \sum_q \frac{1}{\tau(k, q)} = \int \frac{d^3q}{(2\pi)^3} \frac{1}{\tau(k, q)} \quad \vec{q} \equiv \vec{k} - \vec{k}'$$

$$= \underbrace{N}_{\text{ions}} \int \frac{d^3q}{(2\pi)^3} \frac{2\pi}{\hbar} \left(\frac{e^2}{\epsilon_r \epsilon_0 q^2} \right)^2 \delta(E_k - E_{k-q})$$



$$\Rightarrow d^3q = k^2 \sin\theta \, d\theta \, d\phi \, dk$$

$$\frac{1}{T(k)} = \frac{2\pi n}{(2\pi)^3 \hbar} \left(\frac{e^2}{\epsilon_0}\right)^2 \int_0^{2\pi} d\varphi \int_0^\pi d\theta \int_0^\infty dk' \delta(E_k - E_{k'}) \frac{k'^2 \sin\theta}{\epsilon_r^2 q^4}$$

$\frac{dk'}{dE_{k'}} dE_{k'} = \frac{m}{\hbar^2 k'} dE_{k'}$

$$= \frac{2\pi n}{(2\pi)^3 \hbar} \left(\frac{e^2}{\epsilon_0}\right)^2 (2\pi) \int_0^\pi d\theta \left(dE_{k'} \delta(E_k - E_{k'}) \frac{m}{\hbar^2 k'} \frac{k'^2}{\epsilon_r^2 q^4} \sin\theta \right)$$

$$= \frac{2\pi n}{\hbar} \left(\frac{e^2}{2\pi \epsilon_0}\right)^2 \frac{m}{\hbar^2} \int_0^\pi d\theta \frac{k}{\epsilon_r^2 q^4} \sin\theta \quad q = 2k \sin \frac{\theta}{2}$$

$$= \frac{2\pi n}{\hbar} \left(\frac{e^2}{2\pi \epsilon_0}\right)^2 \frac{m}{\hbar^2} \int_0^\pi d\theta \frac{k \sin\theta}{\epsilon_r^2 \hbar k^4 \sin^4 \frac{\theta}{2}} \quad \rightarrow 2 \sin \frac{\theta}{2} \cos \frac{\theta}{2}$$

$$= \frac{2\pi n}{\hbar} \left(\frac{e^2}{2\pi \epsilon_0}\right)^2 \frac{m}{\hbar^2} \frac{1}{8k^3} \int_0^\pi \frac{\cos \frac{\theta}{2} d\theta}{\epsilon_r^2 \sin^3 \frac{\theta}{2}} \quad \begin{aligned} \sin \frac{\theta}{2} &= \eta \\ \frac{1}{2} \cos \frac{\theta}{2} d\theta &= d\eta \end{aligned}$$

$$= \frac{2\pi n m}{\hbar^3 k^3} \left(\frac{e^2}{4\pi \epsilon_0}\right)^2 \int_0^1 \frac{d\eta}{\epsilon_r^2 \eta^3} \quad \begin{cases} E = \frac{\hbar^2 k^2}{2m} \rightarrow k = \frac{\sqrt{2mE}}{\hbar} \\ k^3 = \frac{(2m)^{3/2}}{\hbar^3} E^{3/2} \end{cases}$$

$$= \frac{2\pi n m}{\hbar^3} \frac{\hbar^3 E^{-3/2}}{2^{3/2} m^{3/2}} \left(\frac{e^2}{4\pi \epsilon_0}\right)^2 \int_0^1 \frac{d\eta}{\epsilon_r^2 \eta^3}$$

$$\frac{1}{\tau(E)} = \frac{\pi}{2^{1/2} m^{1/2}} n \left(\frac{e^2}{4\pi\epsilon_0} \right)^2 E^{-3/2} \int_0^1 \frac{d\eta}{\epsilon_r \eta^3}$$

If we assume $\epsilon_r(q) = \epsilon_r = \text{constant} \Rightarrow$

$$\frac{1}{\tau(E)} = \frac{\pi}{2^{1/2} m^{1/2}} n \left(\frac{e^2}{4\pi\epsilon_0 \epsilon_r} \right)^2 E^{-3/2} \underbrace{\int_0^1 \frac{d\eta}{\eta^3}}_{\left. \frac{-1}{2\eta^2} \right|_0^1 = \frac{-1}{2} + \infty !}$$

$$\Rightarrow \frac{1}{\tau(E)} = \infty !$$

So assuming ϵ_r to be constant, doesn't give a meaningful result.

Ex calculate the electron meanfree path and mobility in GaAs doped to $n = 10^{18} \text{ cm}^{-3}$ assuming that the dominant scattering for electrons is due to the ionized impurities.

$$\text{Assume } \int_0^1 \frac{d\eta}{\epsilon_r(q) \eta^3} \approx \frac{1}{\epsilon_{r0}^2} = \frac{1}{100} \quad ; \quad m_e^* = 0.07$$

$$\frac{1}{\tau(E)} = \frac{\pi}{(2m)^{1/2}} n \left(\frac{e^2}{4\pi\epsilon_0} \right)^2 E^{-3/2} \frac{1}{100}$$

What is E ? Most electrons are around Fermi energy - especially at low temperature:

$$E \approx E_f$$

At $T \approx 0\text{K}$, we had: $k_F \approx (3\pi^2 n)^{1/3} = (3\pi^2 \times 10^{18})^{1/3} = 3 \times 10^{16} \text{ cm}^{-1}$

$$E_F = \frac{\hbar^2 k_F^2}{2m_e} \Rightarrow$$

$$\frac{1}{\tau(E)} \approx \frac{\pi \times 10^{18} \times 10^6}{(2 \times 0.07 \times 9.1 \times 10^{-31})^{1/2}} \left(\frac{1.6 \times 10^{-19}}{4\pi \times 8.854 \times 10^{-12}} \right)^2 \times \frac{E_F^{-3/2}}{100}$$

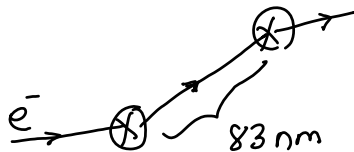
$$\approx 6.14 \times 10^{12} \text{ s}^{-1}$$

The electron scatters 6.14×10^{12} times per second!

The electron mean free path is:

$$l = v\tau \approx v_F \tau(E_F) = 5 \times 10^7 \times \frac{1}{6.14 \times 10^{12}} = 83 \text{ nm}$$

$$v_F = \frac{\hbar k_F}{m_e^*} = 5 \times 10^7 \text{ cm/s}$$



Let's compare with electron wavelength:

$$\lambda = \frac{2\pi}{k} \approx \frac{2\pi}{3 \times 10^6 \text{ cm}^{-1}} \approx 20 \text{ nm}$$

Note: This is comparable to l . So we cannot assume that scattering event is localized upon which the Fermi Golden Rule is based on. We must have $l \gg \lambda$ so the scatterings to be independent.

Mobility calculation:

$$\mu = \frac{e\tau}{m^*}$$

$$= \frac{(1.6 \times 10^{-19})(6.4 \times 10^{12})}{(0.07)(9.1 \times 10^{-31})}$$

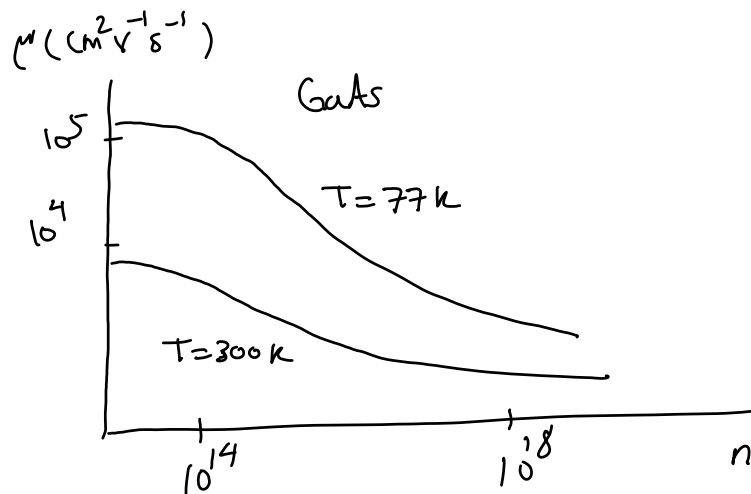
$$\mu = 4100 \text{ cm}^2 \text{V}^{-1} \text{s}^{-1}$$

Conductivity:

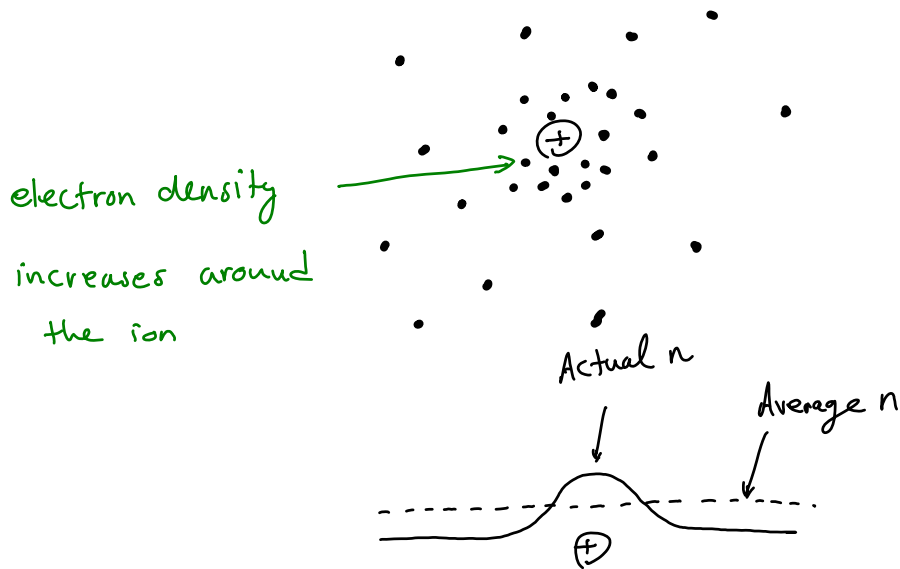
$$\sigma = en\mu$$

$$= (1.6 \times 10^{-19})(10^{18} \text{ cm}^{-3})(4100 \text{ cm}^2 \text{V}^{-1} \text{s}^{-1})$$

$$= 6560 \text{ } \Omega^{-1} \text{cm}^{-1}$$



Screening Effect



We should solve the Poisson equation to find the potential V . Refer to the book for this. But the end result is interesting:

Solve the Poisson's Equation: $\nabla^2 V = -\frac{\rho}{\epsilon_r \epsilon_0} \Rightarrow$

$$V(r) = \frac{-e^2}{4\pi\epsilon_0\epsilon_r} \frac{1}{r} \xrightarrow[\text{screening effect}]{\text{screening effect}} V(r) = \frac{-e^2}{4\pi\epsilon_0\epsilon_r} \frac{e^{-r/r_0}}{r}$$

r_0 = the screening length

$$r_0 = \begin{cases} \frac{1}{q_D} = \left(\frac{\epsilon_r \epsilon_0 k_B T}{n_0 e^2} \right)^{1/2} & \text{Non degenerate semicond.} \\ & \text{Debye screening length} \\ \frac{1}{q_{TF}} = \left(\frac{\epsilon_r \pi^2 \hbar^2}{k_F m e^2} \right)^{1/2} & \text{Degenerate semicond.} \\ & \text{Thomas-Fermi screening length} \end{cases}$$

Fourier transform:

$$v(q) = \frac{-e^2}{\epsilon_0 \epsilon_r} \frac{1}{q^2} \xrightarrow{\text{Screening}} v(q) = \frac{-e^2}{\epsilon_0 \epsilon_r} \frac{1}{(q^2 + \frac{1}{r_0^2})}$$

$$\text{or: } v(q) = \frac{-e^2}{\epsilon_0 \epsilon_r} \frac{1}{q^2} \frac{1}{(1 + \frac{1}{q^2 r_0^2})}$$

$$= \frac{-e^2}{\epsilon_0 \epsilon_r(q)} \frac{1}{q^2}$$

$$q_0 \equiv \frac{1}{r_0}$$

↓

$$\text{As if: } \epsilon_r(q) = \epsilon_{r_0} \left(1 + \frac{1}{q^2 r_0^2}\right) = \epsilon_{r_0} \left(1 + \frac{q_0^2}{q^2}\right)$$

Example

For the Gads of the previous example:

$$q_{TF} = \left(\frac{k_F m e^2}{\epsilon \pi^2 \hbar^2} \right)^{1/2} = 2 \times 10^{16} \text{ cm}^{-1}$$

$$\Rightarrow \frac{1}{q_{TF}} = 5 \text{ nm}$$

Compare it with $\frac{1}{k_F} \approx 3 \text{ nm}$. This is not surprising

that there are close to each other as they are both

a measure of highest spatial frequency that can

be used by electron to screen the coulomb

interaction.

Ionized Impurity scattering - again!

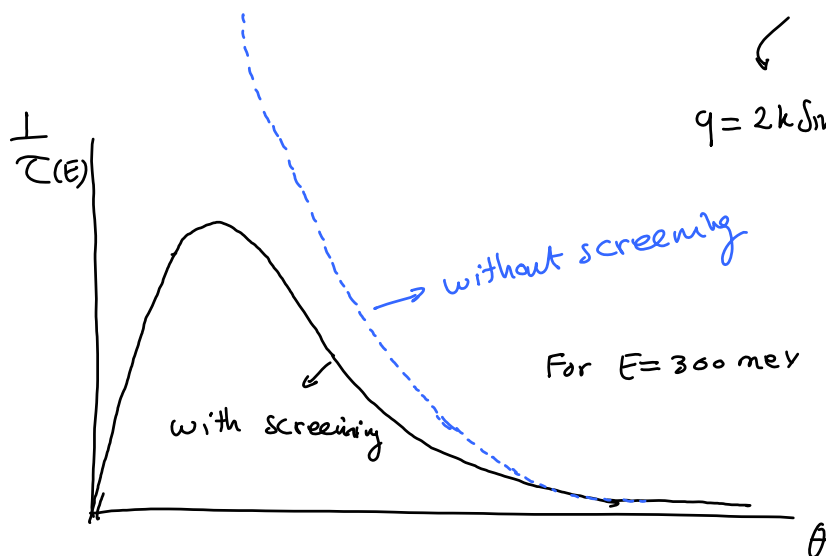
We now include the screening effect to our previous calculation:

$$\frac{1}{\tau(E)} = \frac{\pi}{2^{1/2} m^{1/2}} n \left(\frac{e^2}{4\pi\epsilon_0} \right)^2 E^{-3/2} \int_0^1 \frac{d\eta}{\epsilon_r \eta^3}$$

$$\epsilon_r(q) = \epsilon_{r0} \left(1 + \frac{q_{TF}^2}{q^2} \right)$$

$$= \frac{\pi n}{2^{1/2} m^{1/2}} \left(\frac{e^2}{4\pi\epsilon_0 \epsilon_{r0}} \right)^2 E^{-3/2} \int_0^1 \frac{d\eta}{\eta^3 \left(1 + \frac{q_{TF}^2}{q^2} \right)^2}$$

$$q = 2k \sin \frac{\theta}{2} = 2k\eta$$



$$\frac{1}{\tau(E)} = \frac{\pi}{\sqrt{2}} \left(\frac{e^2}{\epsilon_r \epsilon_0} \right)^2 \frac{E^{-3/2}}{m^{1/2}} \left[\ln \left(1 + \left(\frac{8mE}{\hbar^2 q_0} \right)^2 \right) - \frac{1}{1 + \left(\frac{\hbar^2 q_0}{8mE} \right)^2} \right]$$

$q_0 = q_{TF}$ if degenerate & $q_0 = q_D$ if non-degenerate.

Let's consider two limiting cases of very small and very large electron velocities:

Small initial velocity:

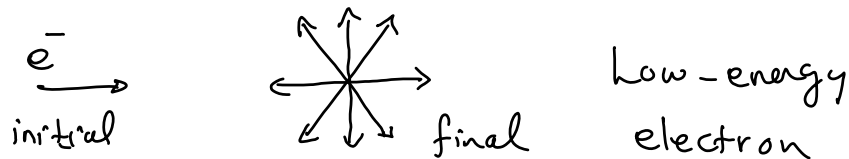
$$k \rightarrow 0 \Rightarrow kr_0 \ll 1.$$

The scattering amplitude $\propto V(q) = \int d^3r V(r) e^{-iq \cdot r}$

$$kr_0 \ll 1 \Rightarrow e^{-iq \cdot r} \sim 1 \quad (\text{since } q = 2k \sin \frac{\theta}{2} \text{ \& } k \rightarrow 0)$$

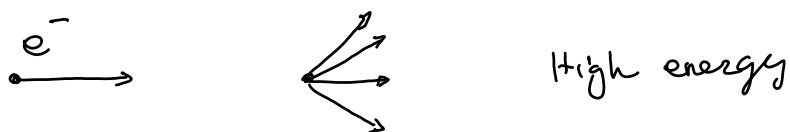
$$V(q) = \int d^3r V(r) = 4\pi \int dr V(r) r^2$$

So it doesn't depend on θ and is **isotropic**:



Large initial velocity:

$k \rightarrow \infty \Rightarrow kr_0 \gg 1$, you can show now it is highly anisotropic:



Stimulated Photon emission

Density of optical modes (states) in 3D:

$$D_3^{opt}(k) = 2 \times \frac{d^3k}{(2\pi)^3} = 2 \times \frac{4\pi k^2 dk}{(2\pi)^3} = \frac{k^2}{\pi^2} dk$$

two orthogonal modes

For electron: $E = \frac{\hbar^2 k^2}{2m}$ For photon: $\omega = ck$ or in a dielectric:

$$\omega = \frac{c}{n_r} k$$

reduced speed

$$D_3^{opt}(\omega) d\omega = \frac{k^2}{\pi^2} dk$$

$$\Rightarrow D_3^{opt}(\omega) = \frac{k^2}{\pi^2} \frac{dk}{d\omega} = \frac{k^2}{\pi^2} \frac{n_r}{c} = \frac{\omega^2 n_r^2}{\pi^2 c^3} \frac{n_r}{c}$$

$$D_3^{opt}(\omega) = \frac{n_r^3}{\pi^2 c^3} \omega^2$$

mode density per unit energy
per unit volume

Light intensity:

Poynting vector $\vec{S} = \vec{E} \times \vec{H}$ (power flux)

$$\text{Average: } |\overline{S}_{av}| = \frac{1}{2} |\vec{E} \times \vec{H}| = \frac{1}{2} |E_0| |H_0|$$

$$= \frac{1}{2} |E_0| \epsilon_0 c |E_0|$$

$$= \frac{1}{2} c \epsilon_0 |E_0|^2$$

$$\text{Note: } \vec{H} = \sqrt{\frac{\epsilon_0}{\mu_0}} \hat{k} \times \vec{E} = \epsilon_0 c \hat{k} \times \vec{E}$$

propagation direction unit vector

$$|S_{av}| = \frac{1}{2} c \epsilon_0 |E_0|^2 \quad \text{Average power flux} \\ \text{(Average light intensity)}$$

Energy density for photon per unit frequency interval in free space is,

$$U(\omega) = \frac{1}{2} \epsilon_0 |E_0|^2$$

Background photon energy density at thermal equilibrium;

$$U(\omega) = D_3^{opt}(\omega) f(\omega) \hbar \omega \\ \downarrow \\ \text{Bose-Einstein dist.}$$

$$U(\omega) = \frac{\omega^2 n_r^3}{\pi^2 c^3} \frac{1}{e^{\hbar\omega/k_B T} - 1} \hbar \omega$$

$$U(\omega) = \frac{\hbar \omega^3 n_r^3}{\pi^2 c^3} \frac{1}{e^{\hbar\omega/k_B T} - 1}$$